

THE VERSATILE ELECTRON RADIATION BELT (VERB) CODE: LONG-TERM SIMULATIONS DURING THE VAN ALLEN PROBES MISSION

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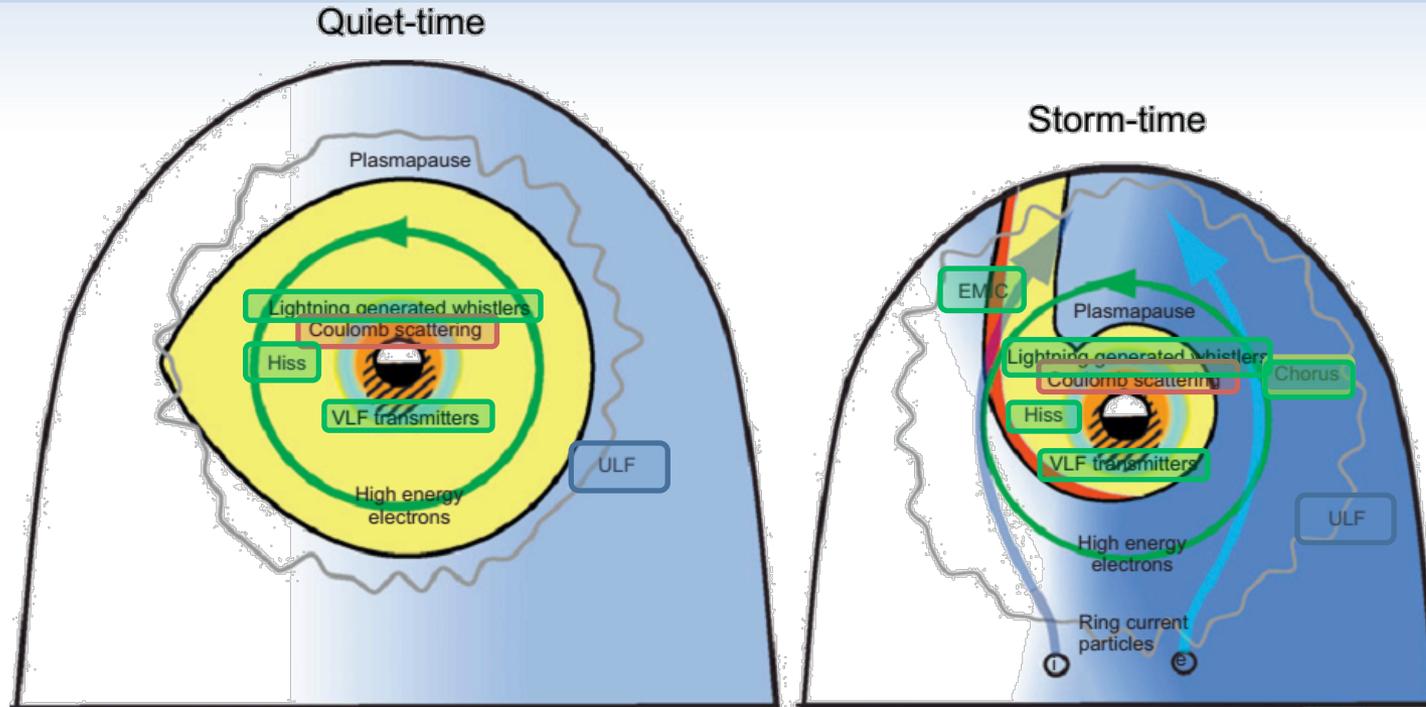
Acknowledgments

N. Aseev, M. Usanova, H. Zhu

Outline

1. The VERB code
 - Long-term simulations
 - Contributions of the EMIC waves
2. VERB 4D
3. Nowcast and forecast
4. Conclusions

Fokker-Planck equation

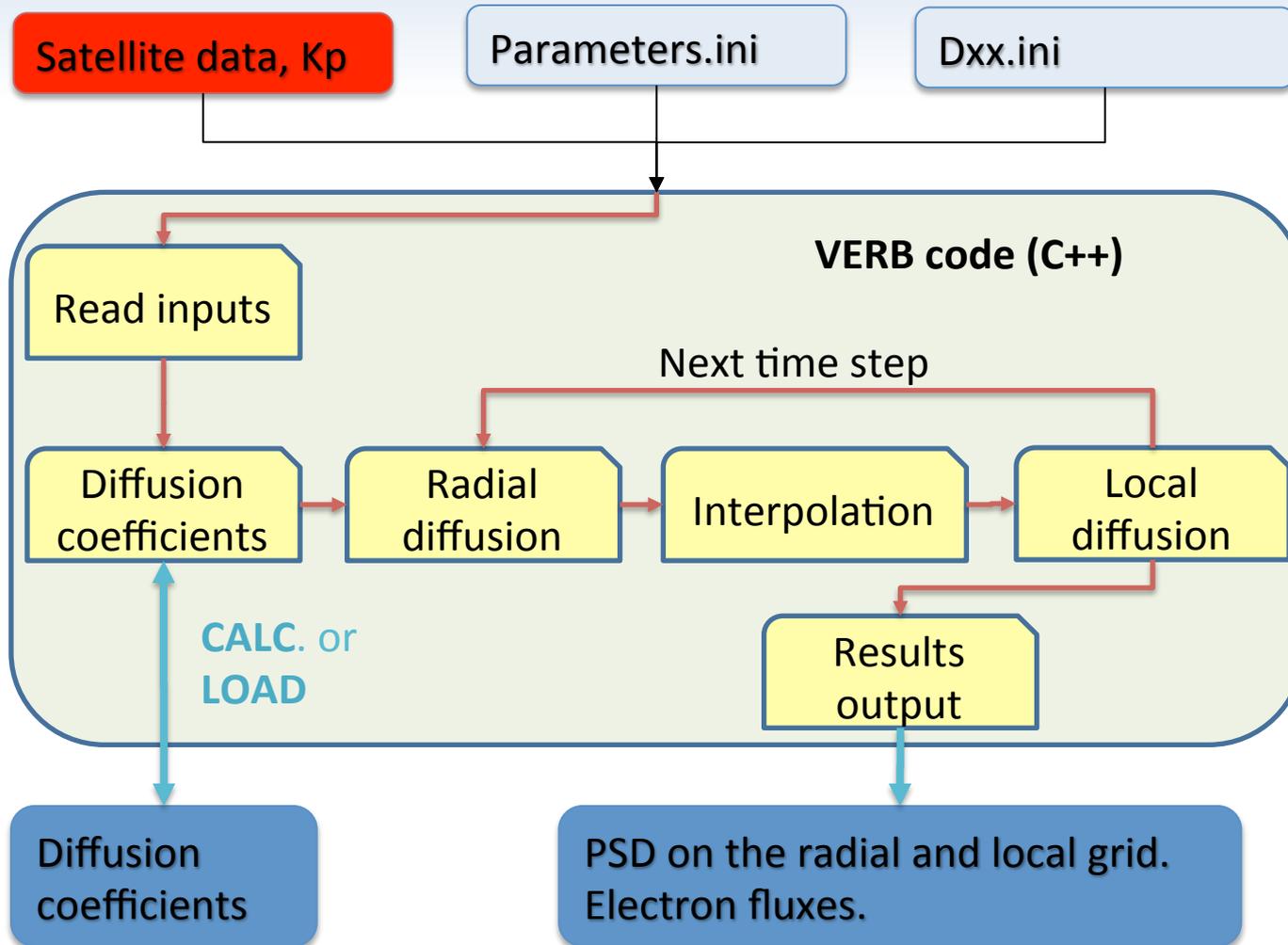


Shprits et al., 2008

A number of different processes determine the evolution of the radiation belt electrons at different levels of geomagnetic activity



VERB code scheme



The VERB code uses an unconditionally stable, implicit scheme and the operator-splitting method to numerically solve the Fokker-Planck equation.

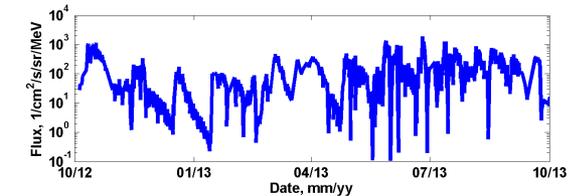
The VERB code

LONG-TERM SIMULATIONS

Long-term simulation with the VERB code and comparison with Van Allen Probes observations

Initial conditions	Steady state	
Boundary conditions	PSD for the whole range of time and energy obtained with MagEIS 0.9 MeV flux variation and averaged spectrum	
Grid	Energy, 0.01 – 10 MeV Pitch-angle, 0.7-89.3 L*, 1 – 5.5 Time step, 1 hour.	101 points. 91 points. 46 points.
Simulation period	• 365 days, Oct 1 st , 2012 – Oct 1 st , 2013	
Losses	Coulomb scattering (loss into atmosphere), Magnetopause shadowing	
Plasmapause	Carpenter and Anderson, 1992	

Boundary conditions



Radial diffusion

$$-9.325 L^{110}$$

Bruutigam and Albert [2000]

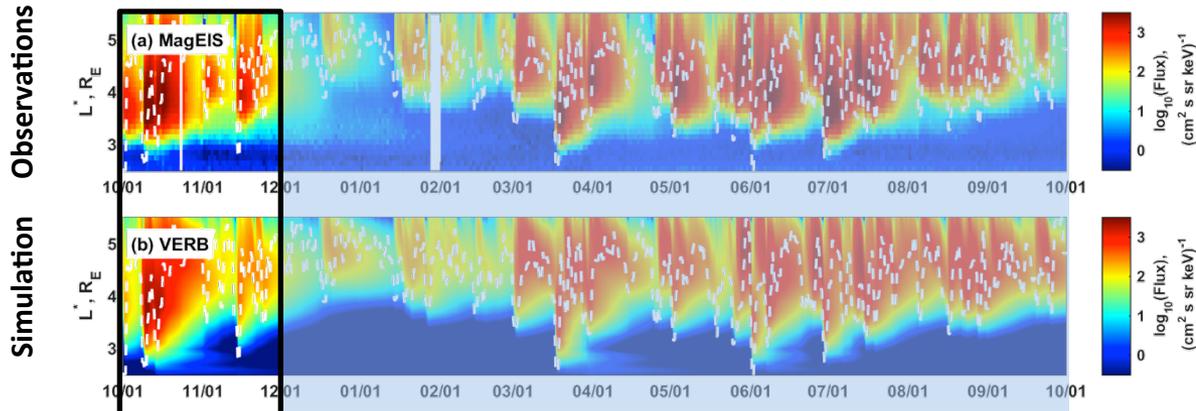
Table of wave parameters

Wave	Amplitude <Kp=2>, [pT]	λ_{\max}	Density model	MLT averaging	Spectral properties and wave normal angle
Chorus day	Kp scaled, <30>	35°	Sheeley et al. 2001	25%	See Subbotin et al. 2011
Chorus night	Kp scaled, <35>	15°		25%	
Plasmaspheric hiss	Kp scaled, <45>	45°	Denton et al. 2004, 2006	60%	Realistic, Orlova et al. 2014 (CRRES)
Lighting, 2xVLF	Kp scaled, <4>, 0.8	45°	Carpenter and Anderson 1992	100%, 4x2.4%	See Subbotin et al. 2011

Long-term VERB code simulation

Oct 2012 – Oct 2013

Flux. Energy 0.9 MeV. L^* vs time plot.



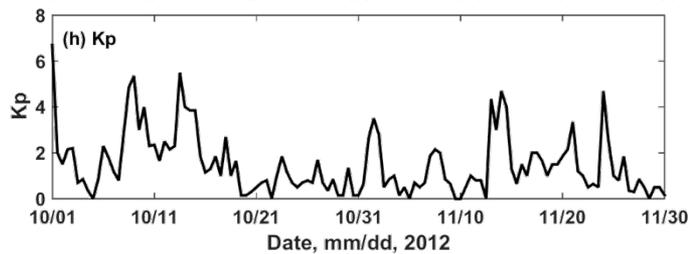
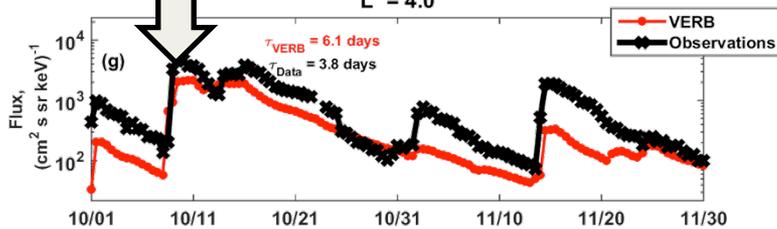
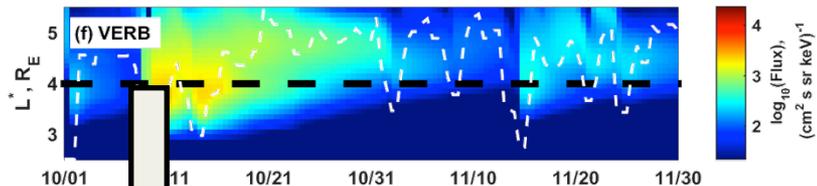
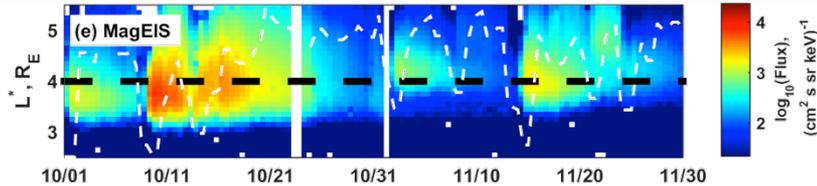
We reproduce the dynamics of relativistic electrons, but not for the ultra-relativistic energies

Decay rates comparison at $L^* = 4$

Oct, 01 – Nov, 30, 2012

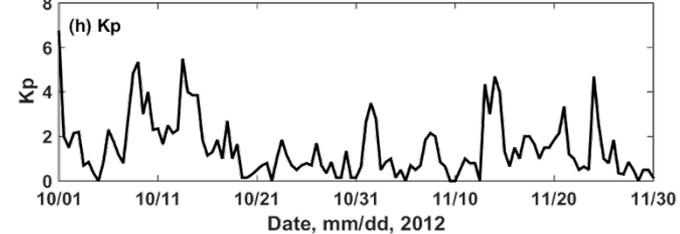
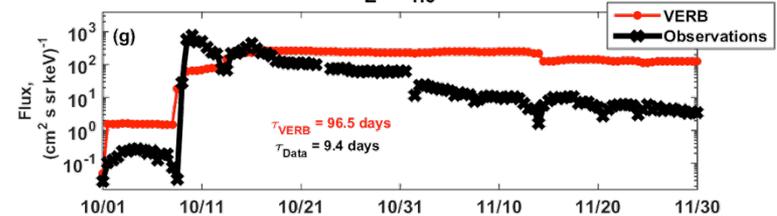
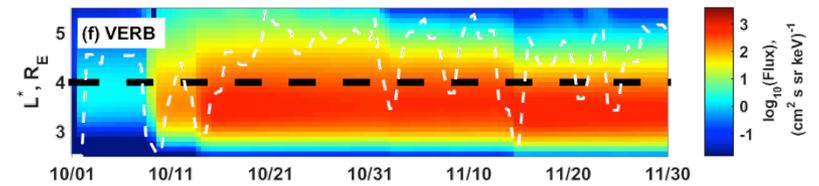
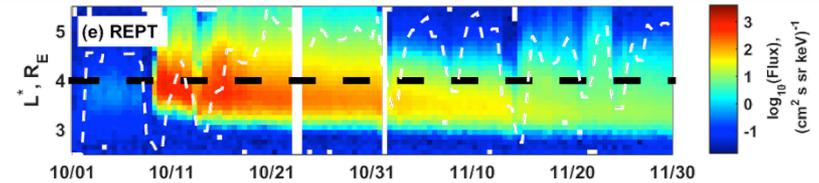
Relativistic electrons. 900 keV

Flux, Energy = 0.9 MeV, $\alpha_{loc} = 85^\circ$



Ultra-Relativistic electrons. 3.6 MeV

Flux, Energy = 3.6 MeV, $\alpha_{loc} = 85^\circ$



$\tau \downarrow VERB = 6.1 \text{ days}, \tau \downarrow Data = 3.8 \text{ days}$

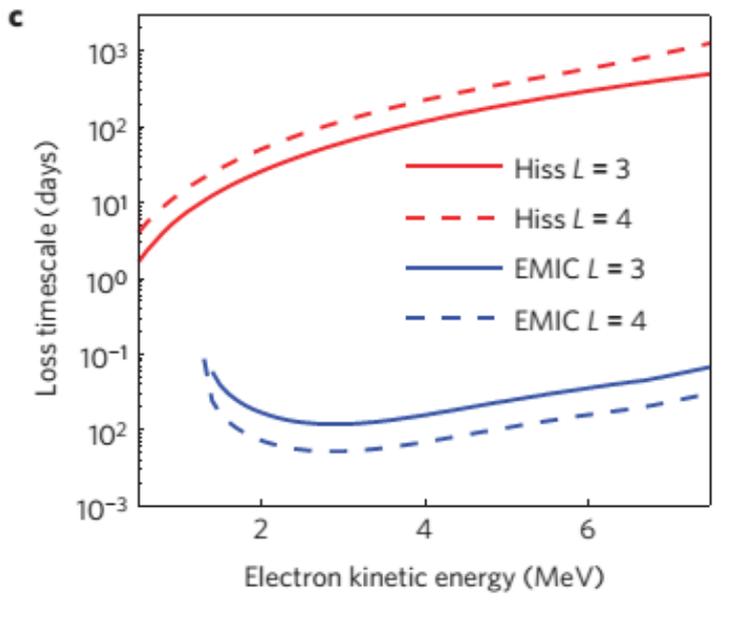
$\tau \downarrow VERB = 96.5 \text{ days}, \tau \downarrow Data = 9.4 \text{ days}$

Observed and modeled decay rates are significantly different for ultra-relativistic electrons

EMIC waves are important!

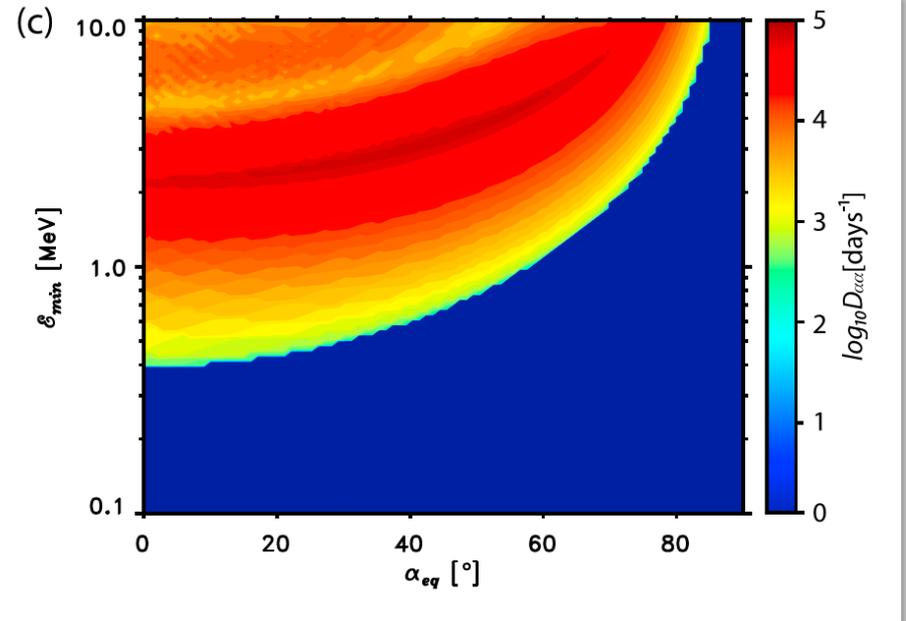
Thorne and Kennel [1971]; Summers and Thorne [2003]; Li et al. [2007] Ukhorskiy et al., [2010] suggested that **EMIC waves may provide efficient scattering mechanism for relativistic electrons.**

The loss time scale for EMIC and hiss waves



Shprits et al., 2014

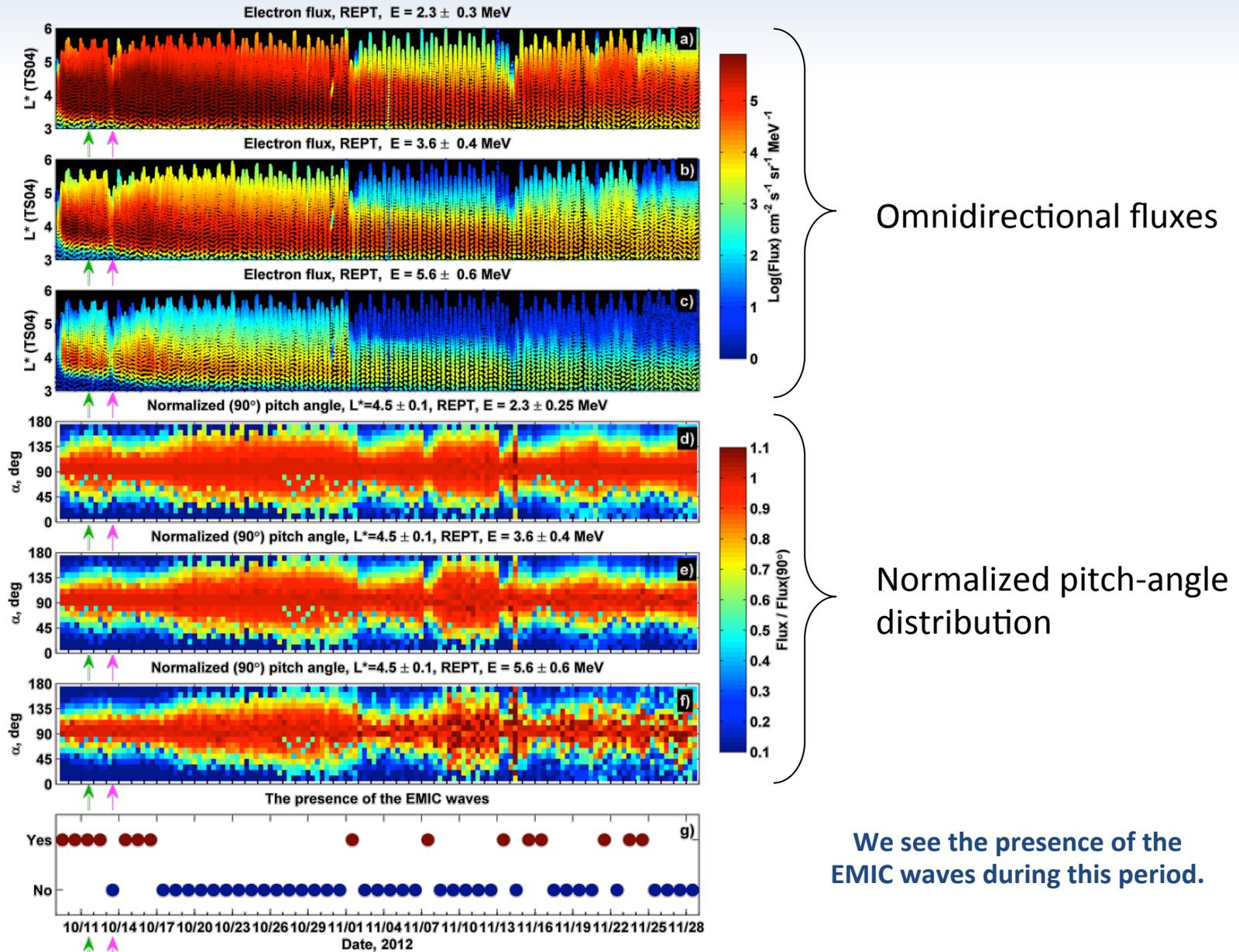
Bounce-averaged pitch-angle diffusion coefficient for the EMIC waves



Ukhorskiy et al. 2010

EMIC waves

Oct, 01 – Nov, 30, 2012



The parametrization of the

EMIC WAVES

EMIC waves parametrizations

$$B_w^2 = 1 \text{ nT}^2 \text{ [Ukhorski et al, 2010; Meredith et al. 2014]}$$

MLT = 25%, Occurrence = 2% [Meredith et al. 2014]

$$K_p \geq 4 \quad S_w \geq 450 \text{ km/s} \quad P \geq 5 \text{ nPa}$$

[Fraser and Nguyen, 2001]

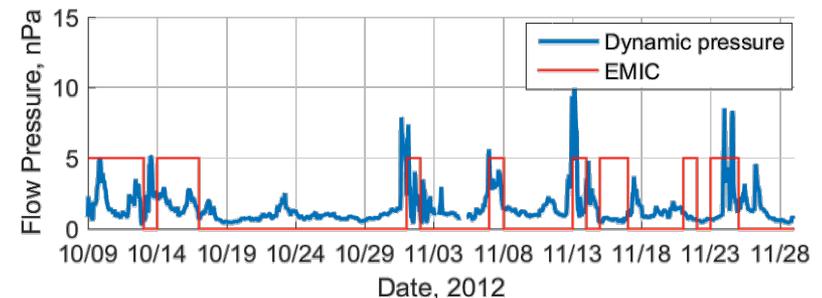
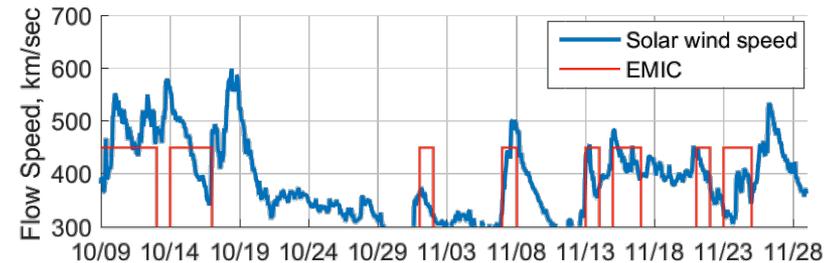
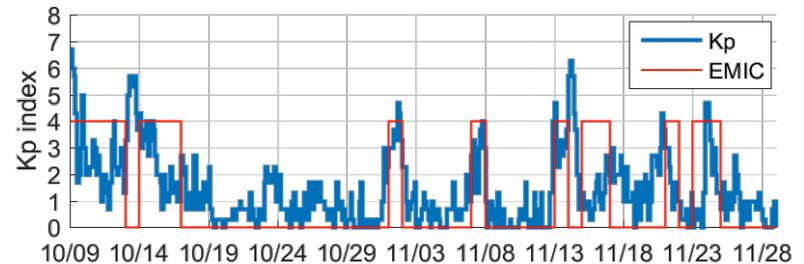
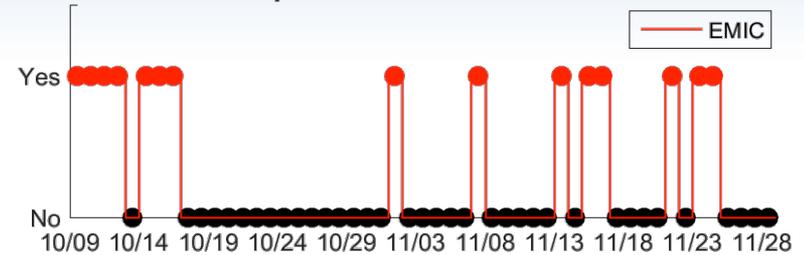
$$\Delta L = L_e - L_{pp}$$

$\Delta L < 0$: inside the plasmopause

$\Delta L > 0$ outside of the plasmopause

$$L = -1 < \Delta L < 4$$

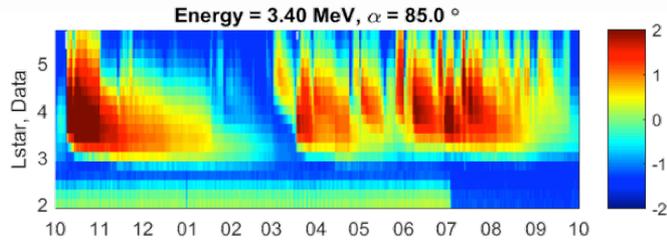
The presence of the EMIC waves



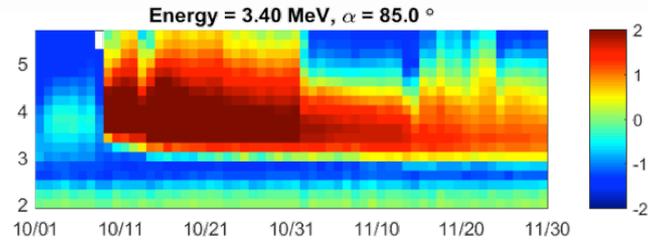
Solar wind velocity ≥ 450 km/s parametrization

Observations

One year



Two months



EMIC:

$$L = -1 < L_{EMIC} - L_{pp} < 4$$

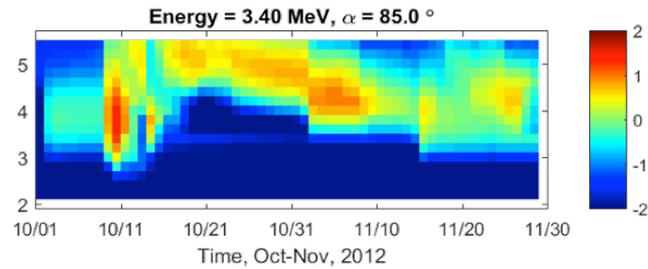
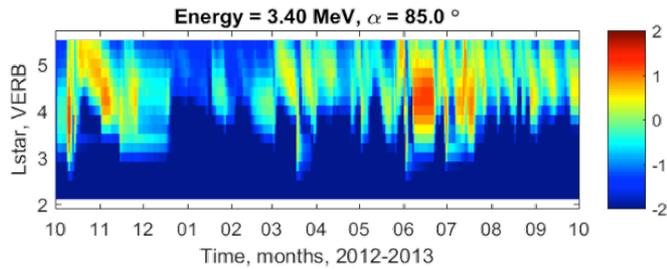
$$Sw \geq 450 \text{ km/s,}$$

$$B_w^2 = 1 \text{ nT}^2$$

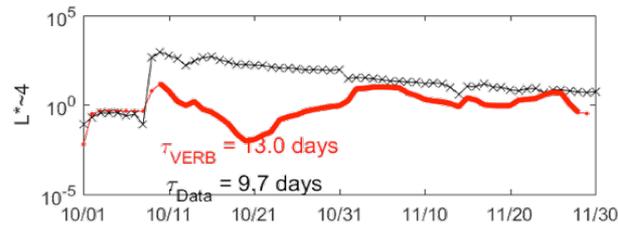
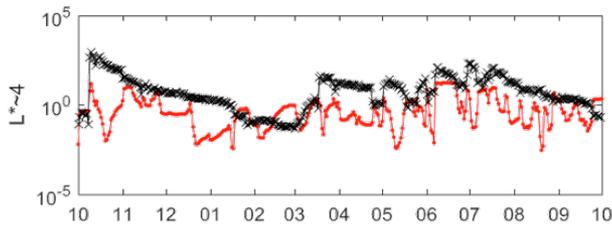
$$\text{MLT} = 25\%,$$

$$\text{Occurrence} = 2\%$$

Simulation



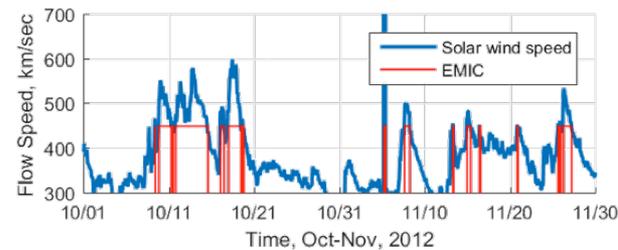
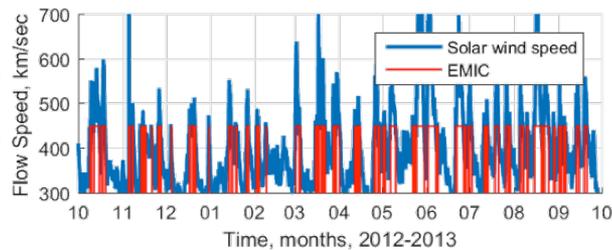
$L^* \sim 4$



— VERB
—x— Data

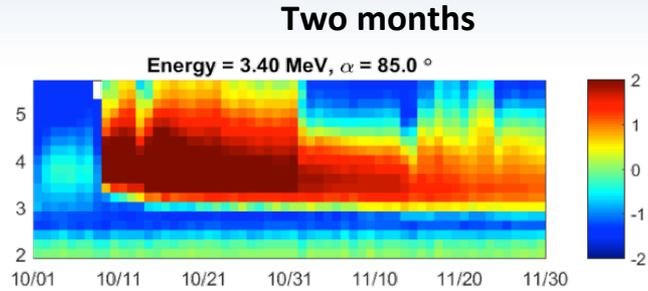
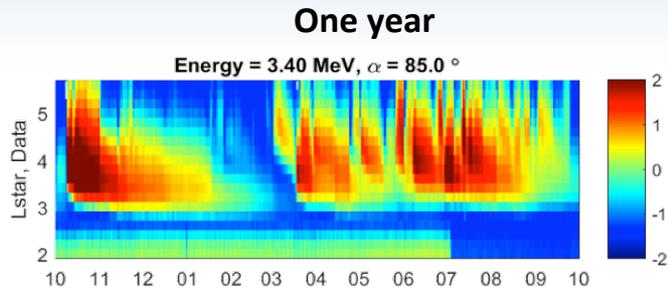
**$\tau \downarrow VERB = 13.0$
days,
 $\tau \downarrow Data = 9.7$
*days***

Index and
EMIC



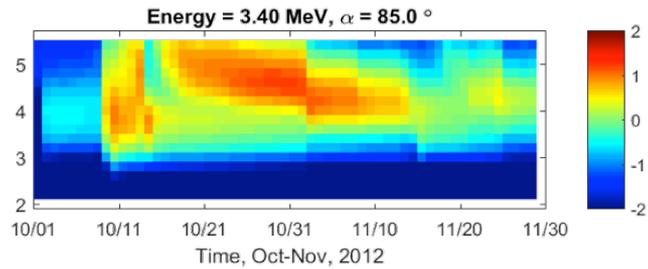
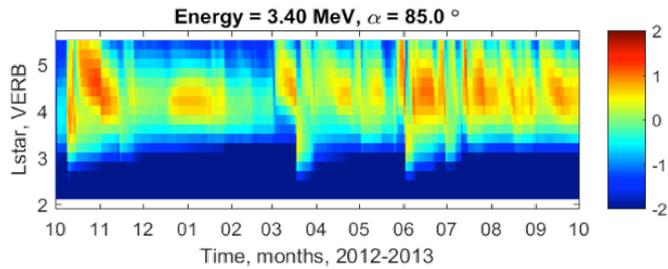
Kp ≥ 4 parametrization

Observations

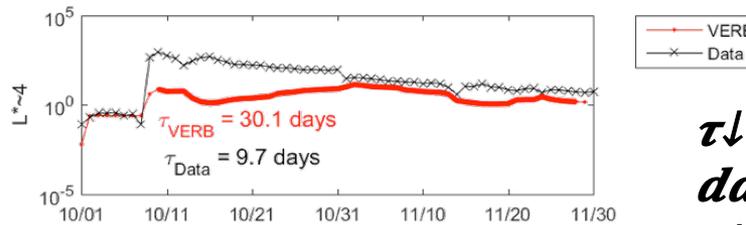
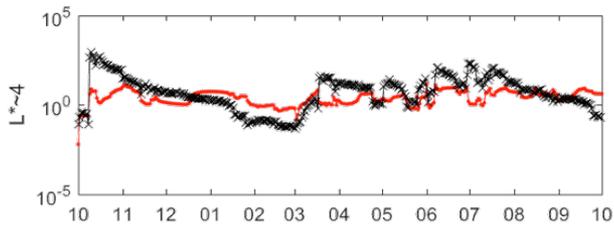


EMIC:
 $L = -1 < L_{EMIC} - L_{pp} < 4$
 $Kp \geq 4$,
 $B_w^2 = 1 \text{ nT}^2$
 $MLT = 25\%$,
 Occurrence = 2%

Simulation



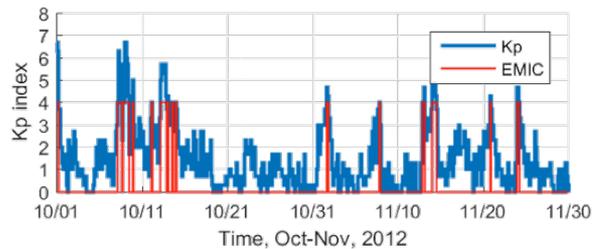
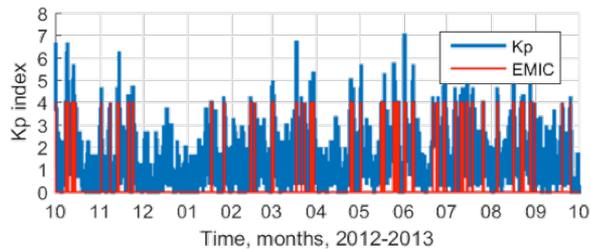
$L^* \sim 4$



— VERB
—x— Data

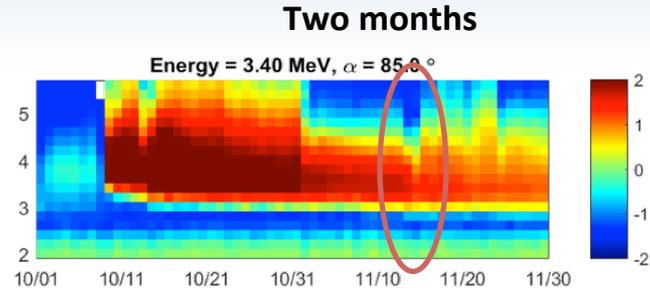
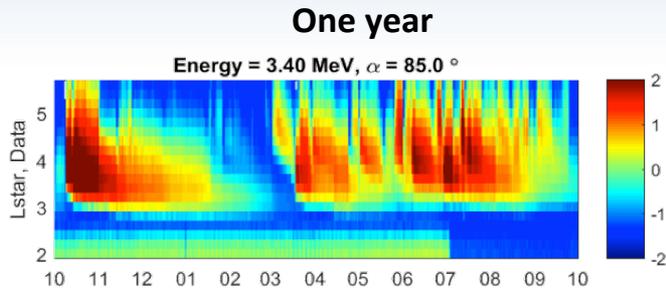
$\tau \downarrow VERB = 30.1$
days,
 $\tau \downarrow Data = 9.7$
days

Index and
EMIC



Solar wind pressure ≥ 5 nT parametrization

Observations



EMIC:

$L = -1 < L_{EMIC} - L_{pp} < 4$

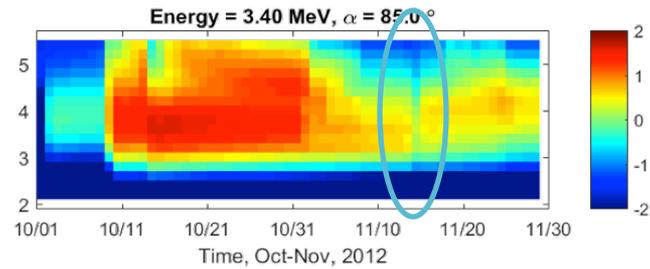
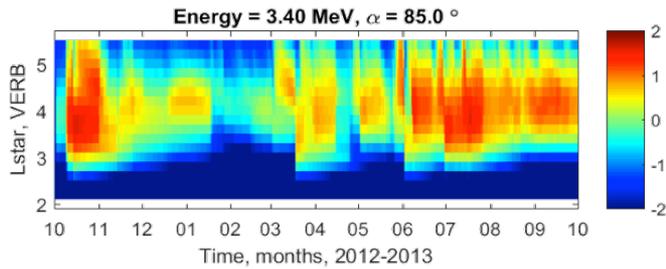
$P \geq 5$ nT,

$B_w^2 = 1$ nT²

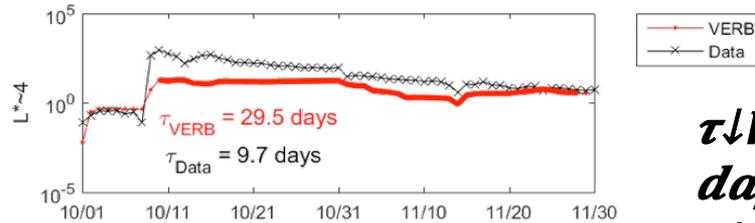
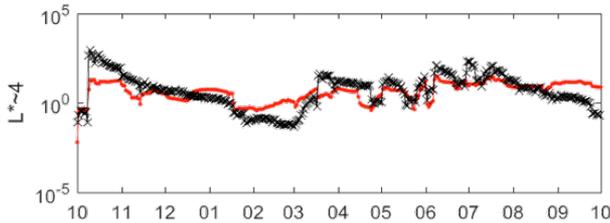
MLT = 25%,

Occurrence = 2%

Simulation



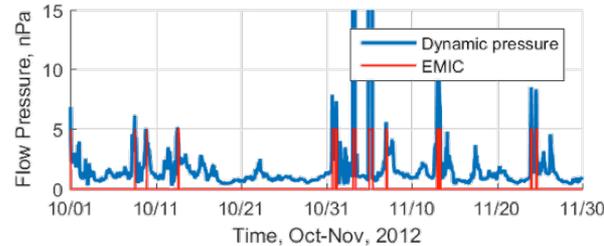
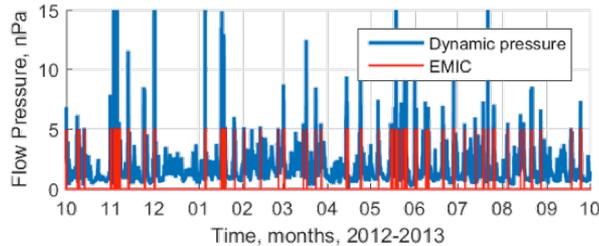
$L^* \sim 4$



$\tau \downarrow VERB = 29.5$
days,

$\tau \downarrow Data = 9.7$
days

Index and
EMIC

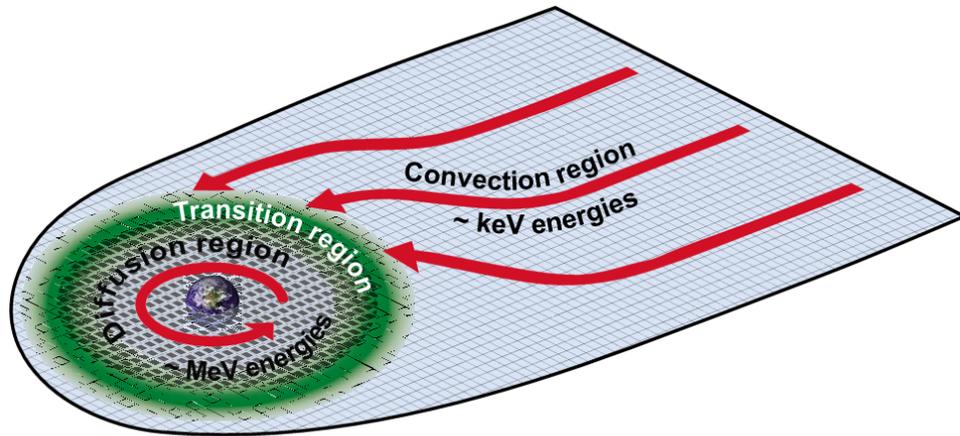


The best agreement
between simulation
results and observations

The new version

VERB 4D CODE

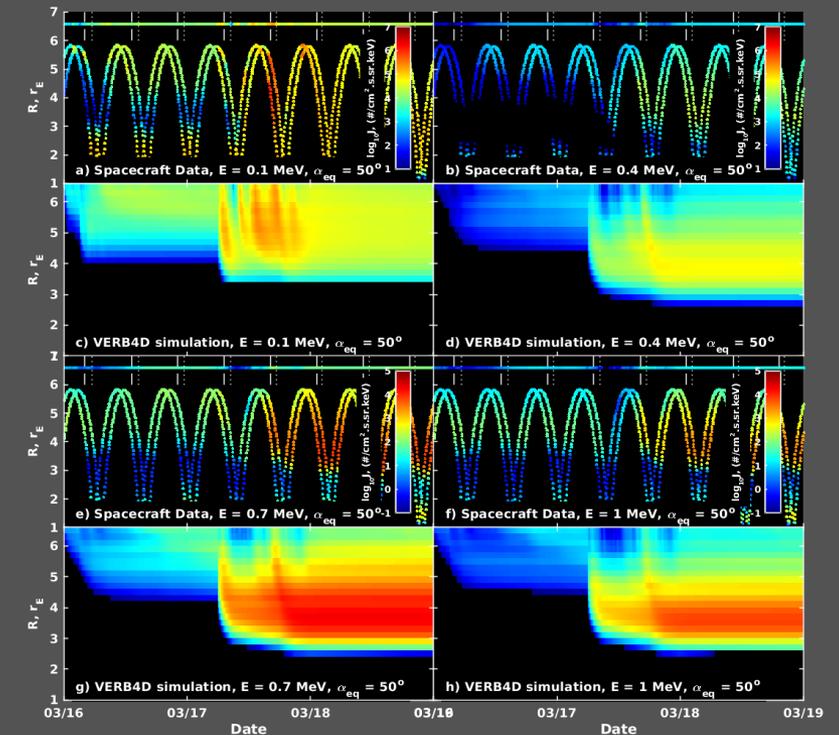
Convective and Diffusive simulations of March 17, 2013 storm



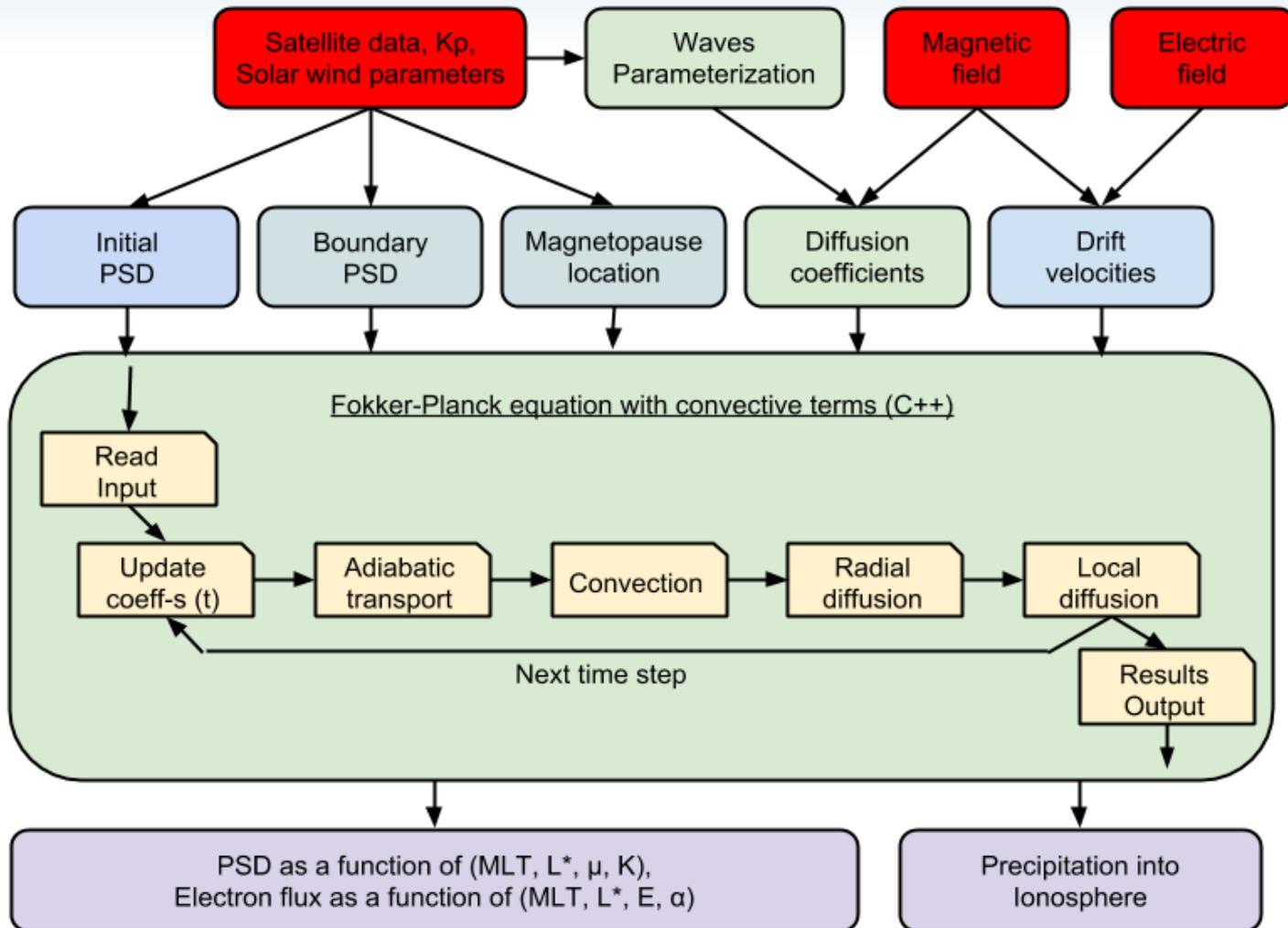
During the March 2013 storm energetic relativistic and ultra-relativistic electrons show very different dynamics. The comparison with observations at various energies can help validate the code and reveal if the dominant physical mechanisms operate at a wide range of energies. Simulations show similar dynamics as observations.



$$\frac{\partial f}{\partial V} - f/\tau$$



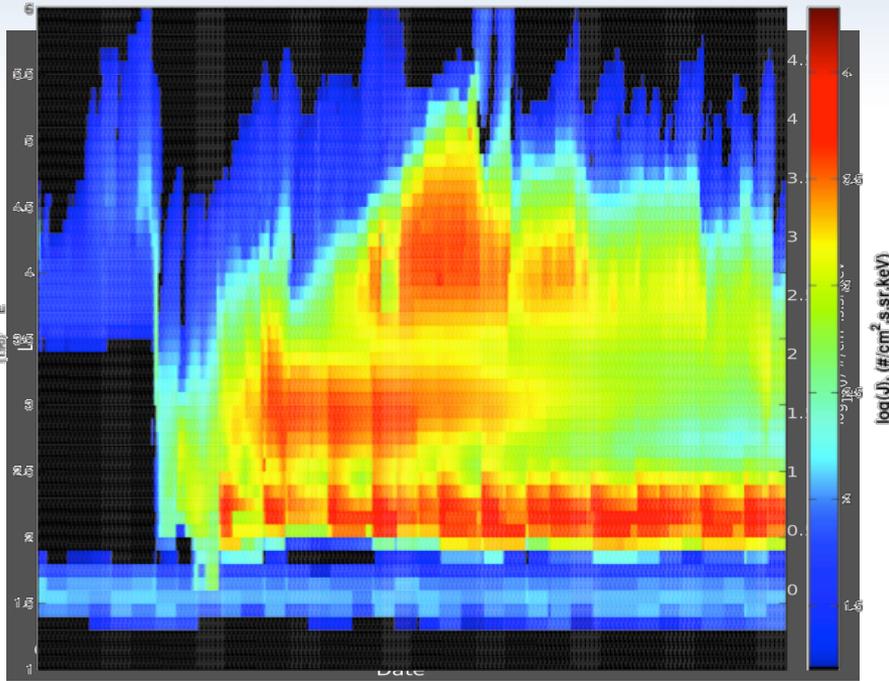
VERB 4D code scheme



Radiation belts

NOWCAST AND FORECAST

Data assimilation and GREEP

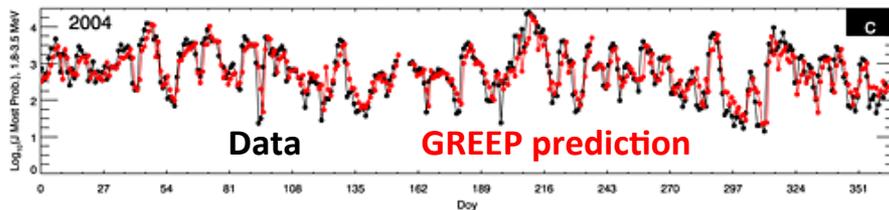


- We can reconstruct the PSD by combining the observations with the VERB code through a split-operator Kalman filter technique.

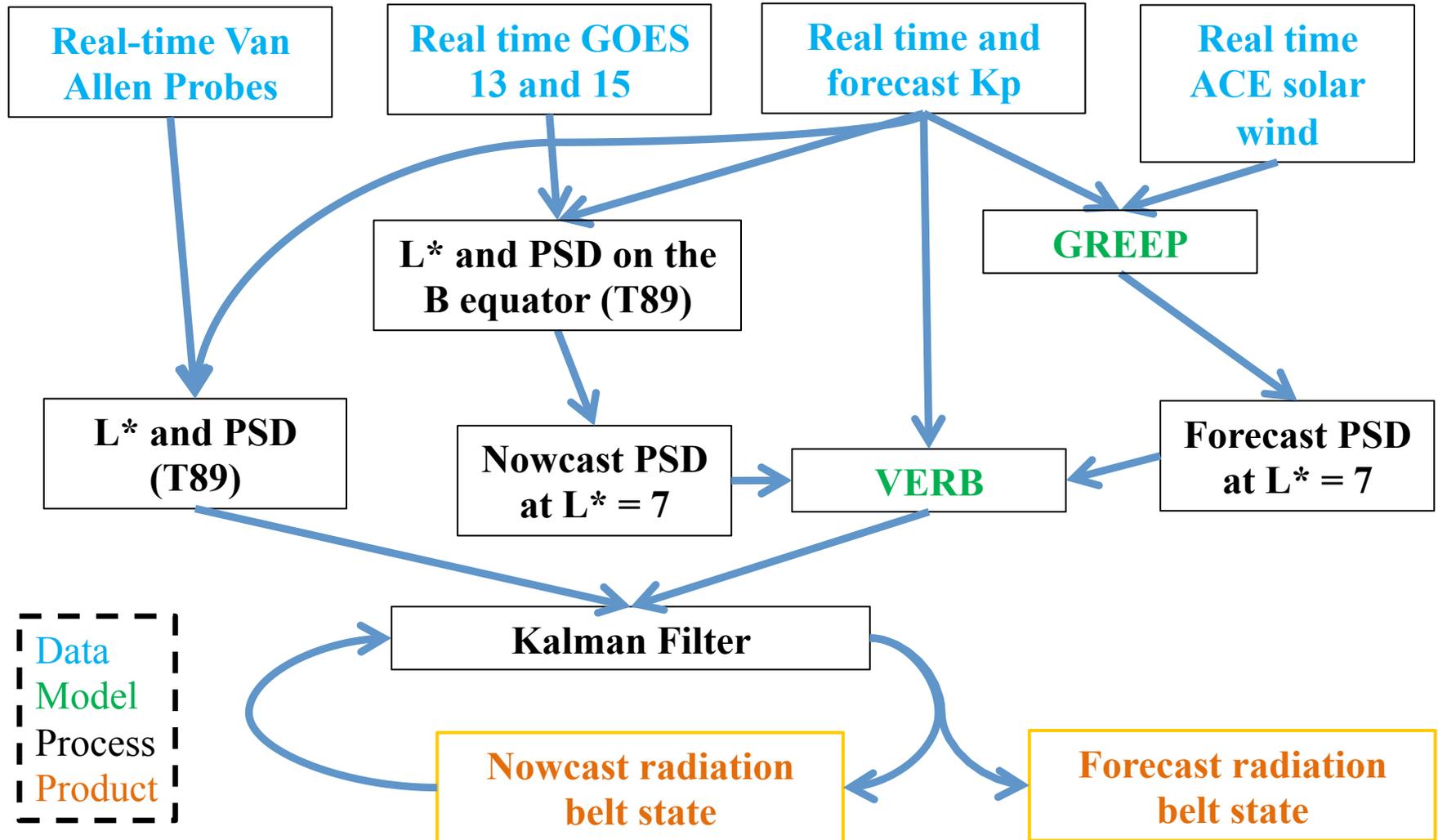
Kellerman et al. (2014), *Three-dimensional data assimilation and reanalysis of radiation belt electrons: Observations of a four-zone structure using five spacecraft and the VERB code*, JGR

- **Geosynchronous Radiation belt Electron Empirical Prediction model (GREEP)** is based on the solar wind velocity and density conditions and can predict the daily averaged outer radiation belt electron fluxes .

Kellerman et al. (2013), *A Geosynchronous Radiation-belt Electron Empirical Prediction (GREEP) model*, Space Weather

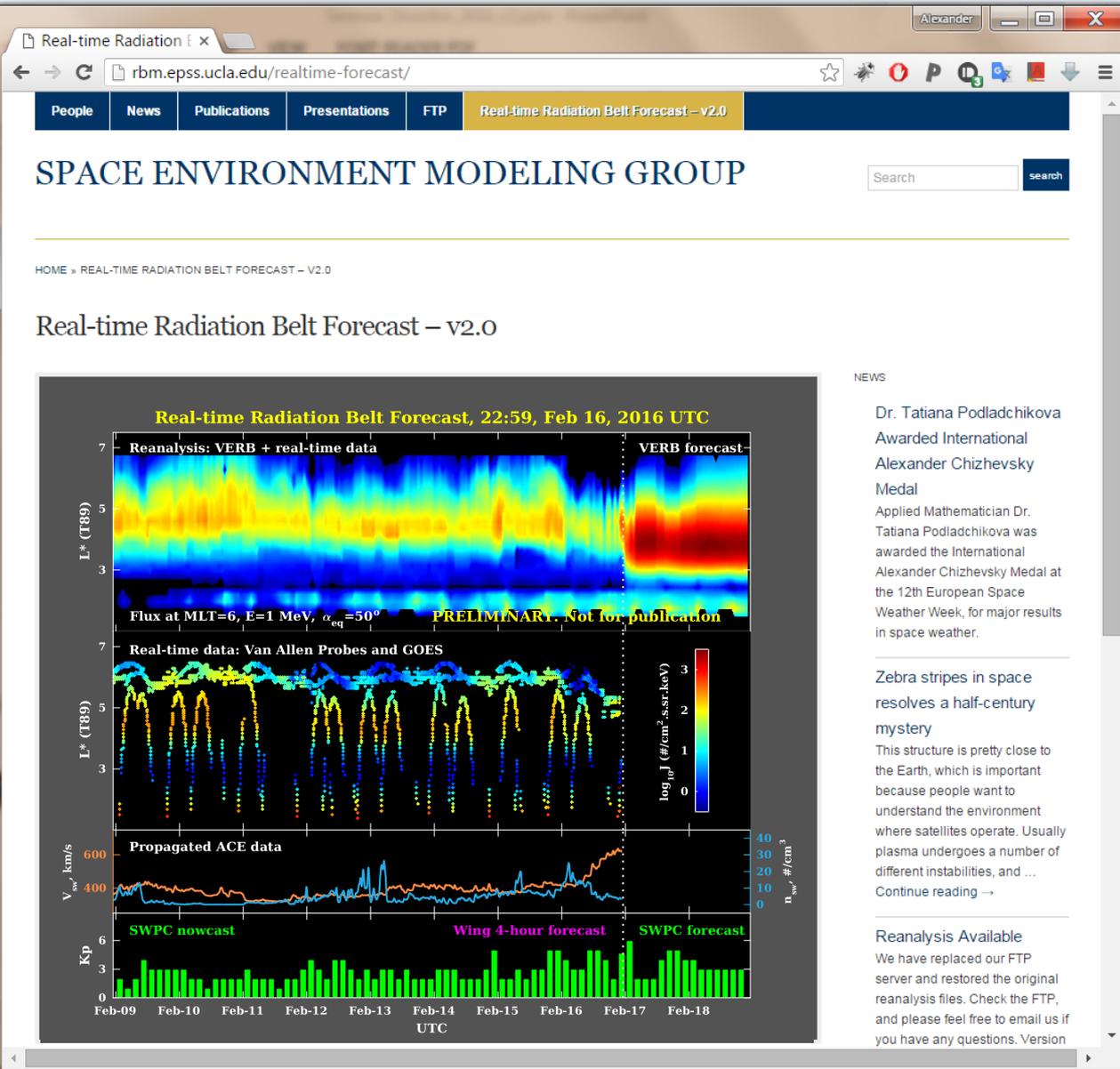


Radiation Belt Forecast Framework



Forecast

Real-time forecast of the
electron radiation belt dynamics
is available on
<http://rbm.epss.ucla.edu>



NEWS

Dr. Tatiana Podladchikova
Awarded International
Alexander Chizhevsky
Medal

Applied Mathematician Dr. Tatiana Podladchikova was awarded the International Alexander Chizhevsky Medal at the 12th European Space Weather Week, for major results in space weather.

Zebra stripes in space
resolves a half-century
mystery

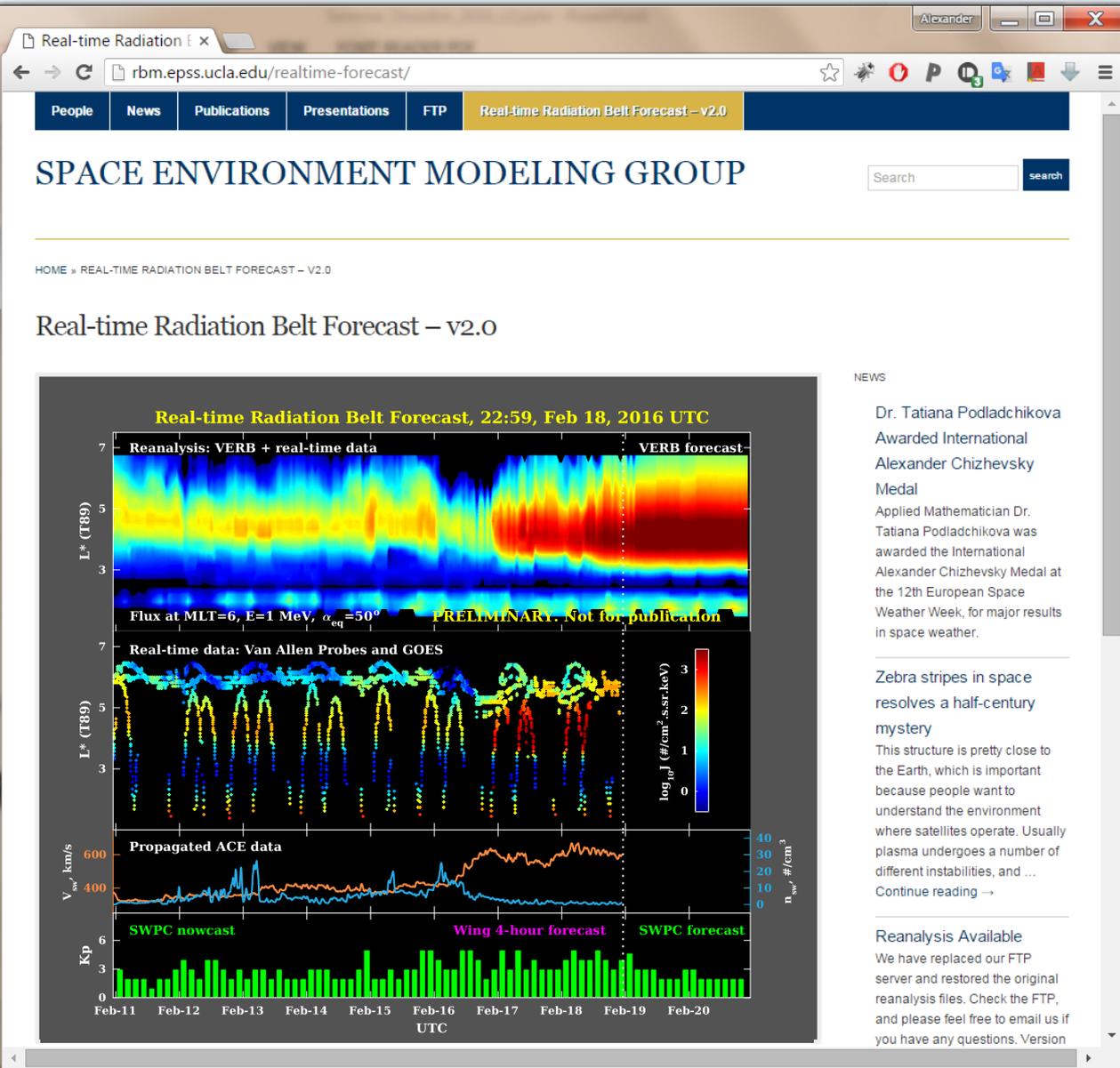
This structure is pretty close to the Earth, which is important because people want to understand the environment where satellites operate. Usually plasma undergoes a number of different instabilities, and ...
Continue reading →

Reanalysis Available

We have replaced our FTP server and restored the original reanalysis files. Check the FTP, and please feel free to email us if you have any questions. Version

Forecast

Real-time forecast of the electron radiation belt dynamics is available on <http://rbm.epss.ucla.edu>



NEWS

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Conclusions

- The Versatile Electron Radiation Belt (VERB) code is capable of solving the Fokker-Planck equation, accounting for radial diffusion and local scattering.
- Comparison between long-term VERB code simulation and the observations shows good agreement for the relativistic (~ 1 MeV) electrons. However, the fluxes for the ultra-relativistic energies (>3 MeV) are overestimated.
- Most likely, EMIC waves are necessary additional losses. The simulation with parametrizes EMIC waves by solar pressure provides better results.
- The new VERB-4D code includes convection processes and uses new invariant grid to calculate radial and local diffusion. Simulations with the 4D code will allow to compare simulations at various MLT with multi point observations provided by Van Allen Probes, THEMIS, Cluster II, MMS, and other missions.
- The reanalysis provides a global reconstruction of the state and evolution of the radiation belts, as compared to sparse spacecraft observations.
- Data assimilation is currently applied to conduct operational forecasting of the electrons radiation belt for 1-2 days into the future.

THANK YOU